

Study of Moon and Sun impacts on Earth's state climate

Abstract

The threat of climate change represents a significant challenge for a considerable proportion of the global population, due to the potential for disasters that it poses. Consequently, it is imperative to conduct comprehensive research on this subject. The objective of this study is to examine the influence of the Moon and Sun on atmospheric tides and their subsequent impact on temperature, wind distribution and precipitation. A model was developed and verified through satellite observations to assess its predictive capacity for climatic phenomena. The findings indicate that the heights of the moon and sun give rise to tidal phenomena that exert a significant climatic influence, which should not be disregarded. Prioritising the study of these phenomena will facilitate a deeper comprehension of the distribution of atmospheric high tides on Earth, which in turn will enable a more accurate prediction of wind patterns, temperatures and precipitation. This will in turn lead to the prevention of climatic disasters.

Keywords: Climate change, Atmospheric tides, Celestial body heights, Precipitation, Winds, Temperatures

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LIST OF SYMBOLS

Δu_c : ----- **Variation vitesse critique**

k_c : ----- **Critical wavenumber**

h_{fluidT} : ----- **Total fluid tide height**

g : ----- $9,81m/s^2$

G : ----- **Gravitation constant**

u_i : ----- **Layer i speed**

χ_i : ----- (\vec{OP}_i, \vec{OM}_i)

ρ_i : ----- **Layer i density**

γ : ----- **Surface tension coefficient**

k : ----- **Angular wave number**

R : ----- **Earth radius**

Δ : ----- **The distance between the centre of the Earth and that of an extraterrestrial object is defined as the Earth-aster distance.**

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I. INTRODUCTION

The phenomenon of climate change is a consequence of alterations in the composition of the atmosphere. This disruption to the atmospheric equilibrium is manifested in the form of an increase in the mean global temperature, which in turn gives rise to modifications in the physical, chemical and biological characteristics of the Earth system [1]. The environmental impacts are numerous, significant and increasingly frequent, including droughts, glacial and sea ice melting, rising sea levels and an increase in the frequency of tropical storms. These effects are felt by the entire global population and have a significant impact on biodiversity. The Intergovernmental Panel on Climate Change (IPCC) has concluded that global warming is a real phenomenon and that human activity is responsible for it, through the emission of greenhouse gases (GHGs). Furthermore, over the last century, the average temperature at the Earth's surface has risen by 0.74°C , whereas it was predicted that the average temperature would only increase by 0.6°C (IPCC, 2008). Projections for the year 2100 are even more disconcerting, with an anticipated increase in the average temperature of between 2 and 4.5°C (IPCC, 2008). Greenhouse gases are naturally present in the atmosphere. These gases form a layer around the Earth, which enables it to retain heat energy. This phenomenon is known as the greenhouse effect. Solar radiation heats the Earth, which then emits some of its heat back into space. The greenhouse gases present in the atmosphere serve to trap a portion of the heat emitted by the Earth, preventing its escape back into space. This phenomenon enables the maintenance of average planetary temperatures of 15°C . In the absence of this phenomenon, the temperature would be approximately -18°C , which would not support the existence of life as we know it (Environment Canada, 2009). The rate of ice melting has increased markedly over the past 15 years. The extensive melting of ice results in an increase in the mass of water, which in turn causes a rise in sea levels. The projected increase in sea level is estimated to range between 9 and 88 centimetres between 1990 and 2100. This situation will have deleterious consequences for low-lying coastal areas, which are home to a significant proportion of the global population. In 2008, climate-related disasters resulted in the displacement of 20 million people (IOCHA and IDMC, 2009).

Furthermore, the availability of drinking water is likely to be impacted. As an illustration, an increase in temperature could result in a significant reduction in the water level of the Great Lakes, due to enhanced evaporation. In addition to being a principal tributary of the St. Lawrence River, the Great Lakes represent the world's largest source of fresh water. A reduction in the quan-

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tivity of drinking water could have significant implications, potentially leading to conflict between nations (Environment Canada, 2002). Furthermore, the quality of lake water will require monitoring, as the proliferation of blue-green algae could intensify in the context of climate change. The phenomenon of climate change is already resulting in disruption to precipitation patterns. Such upheavals will be accompanied by an increase in the frequency and intensity of extreme weather events, including droughts, floods, heatwaves, heavy and abundant precipitation, and tornadoes. Such meteorological occurrences, frequently the catalyst for disasters, may well become more prevalent in the future (Environment Canada, 2008). In Quebec, projections indicate a potential increase in precipitation of up to 25 percent in the spring and 32 percent in the winter by 2050 (Bourque, Simonet, 2007). Notwithstanding these projections, the rate of snow cover in our winters remains uncertain, as precipitation may also manifest as rain. The majority of scientists now acknowledge that climate change will have significant implications for human health. In Canada, a number of health impacts have been identified, including an increase in respiratory diseases such as asthma, skin cancers, malaise and deaths linked to intense heat (Health Canada, 2008). Furthermore, climate change can exacerbate the phenomenon of heat islands in urban areas (Giguère, 2009), which has a range of health impacts, from heat cramps to heat syncope, and from heat exhaustion to heat stroke (Denis, 2010). Furthermore, an increase in vector-borne diseases is to be anticipated. With regard to the agricultural sector, climate change may exert a dual impact, both positive and negative. In general, an increase in average temperatures and a lengthening of the growing season should result in an increase in crop yields. Similarly, these changes should facilitate the production of crops that are adapted to higher temperatures (Bélanger, 2002). Conversely, the winter protection provided by snow cover for perennial crops may be adversely affected by milder winters with less snow. Furthermore, milder autumns may result in suboptimal conditions for hardening and potentially exacerbate damage to forage plants (Bélanger, 2002). The risk of invasion by insect pests could increase, and the distribution of species could change over the next few years as a result of more favourable climatic conditions (Roy, 2002). For example, observations have confirmed an increased likelihood of invasion by the western rootworm and the leafroller. For example, observations have confirmed that the western rootworm and the oriental fruit moth, which were previously present in southern Quebec, are now exhibiting a northward migration. The occurrence of specific insect species is now being documented over a more extended period, necessitating the initiation of surveillance activities at an earlier stage in the season. The expansion of weed habitats is also likely to occur as a result of the favourable conditions brought about by cli-

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mate change. Furthermore, some studies have indicated that weeds may possess greater adaptive capacity to climate change than crops (Ziska, undated). It is imperative that the agricultural sector adapt to the new conditions brought about by climate change. It is evident that the agricultural sector will face significant challenges in the coming years. However, there are opportunities for farmers to adapt to climate change by introducing new crop varieties or new types of production, ensuring better soil protection and improved water conditions. The earlier agricultural producers are made aware of the necessary adaptations to be made in order to mitigate the negative impact of climate change on their farms, the more effectively they will be able to implement these changes. Indeed, the implementation of measures to combat climate change at this juncture is more advantageous than the requisite adaptations that will be necessitated by the impacts. The advancement of scientific research has led to a notable evolution in our understanding of the climate. In light of the knowledge acquired since the signing of the Kyoto Protocol in 1997, climate scientists have posited that countries' targets will have to be revised upwards. A temperature rise in excess of 2°C could precipitate significant climate disruption on a global scale, with the potential for major humanitarian and political crises to ensue. In order to take effective action against global warming, it is essential to ascertain whether greenhouse gases are indeed the primary cause. In order to address this question, this study will also investigate the impact of the position of stars such as the Sun and the Moon on atmospheric tidal phenomena and climate consequences. This will be done with a view to comparing the two models and deducing which of the two points of view has the greatest real impact on climate change.

II. METHODS

In the local reference frame, only the zenith angle and the height can be accurately measured to determine the latitude and local time. The objective is to derive the formulae that enable the height of a star to be related to values of position and time. [2] The celestial sphere, or sphere of fixed points, is linked to the equatorial coordinate system (right ascension and declination). The observer is located in two reference frames related to their position: the local reference frame (azimuth and height) and the hourly reference frame (hourly angle and declination). The observer is situated within two distinct reference frames, both of which are related to their position[3]. The first of these is the local reference frame, which encompasses azimuth and height. The second is the time reference frame, which incorporates time angle and declination [4]. These two reference

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$$\sin z \cdot \sin a = \cos \delta \cdot \sin H \tag{6}$$

$$\sin z \cdot \cos a = -\cos \varphi \cdot \sin \delta + \sin \varphi \cdot \cos \delta \cdot \cos H \tag{7}$$

The complement of the height is found by using the latitude of the location, denoted by φ , and the zenith distance, denoted by z [6]. In the figure opposite, all these coordinates are found in the miracle triangle AZP:

- angle APZ hour angle
- angle PZA complement of the azimuth
- arc PA or angle PCA: declination supplement
- arc ZA or angle ZCA: complement of height
- arc PZ or angle PCZ: latitude complement

The formula for the PZA triangle is as follows:

$$\cos ZA = \cos PA \cdot \cos PZ + \sin PA \cdot \sin PZ \cdot \cos H \tag{8}$$

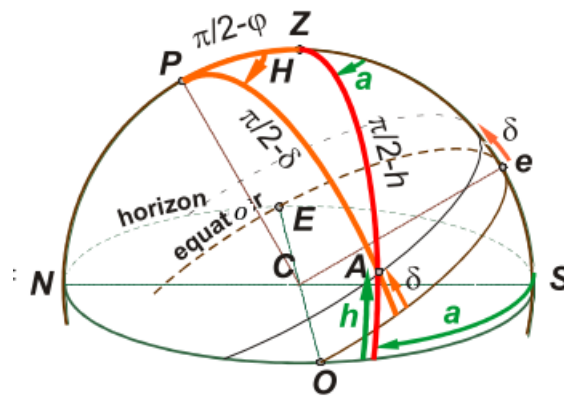


FIG. 2. PZA Triangular

$$\sin h = \sin \delta \cdot \sin \varphi + \cos \delta \cdot \cos \varphi \cdot \cos H \tag{9}$$

$$\cos H = \frac{\sin h - \sin \delta \cdot \sin \varphi}{\cos \delta \cdot \cos \varphi} \tag{10}$$

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Using the sine formula, we obtain the azimuth:

$$\frac{\sin a}{\cos \delta} = \frac{\sin H}{\cos h} \tag{11}$$

With

$$\sin a = \frac{\sin H \cdot \cos \delta}{\cos h} \tag{12}$$

The heights of the Moon and the Sun are not fixed throughout the year; rather, they vary continuously each day, which generates tidal phenomena on the fluids (liquids (water reservoirs), gases (atmosphere)) on the surface of the Earth[7].

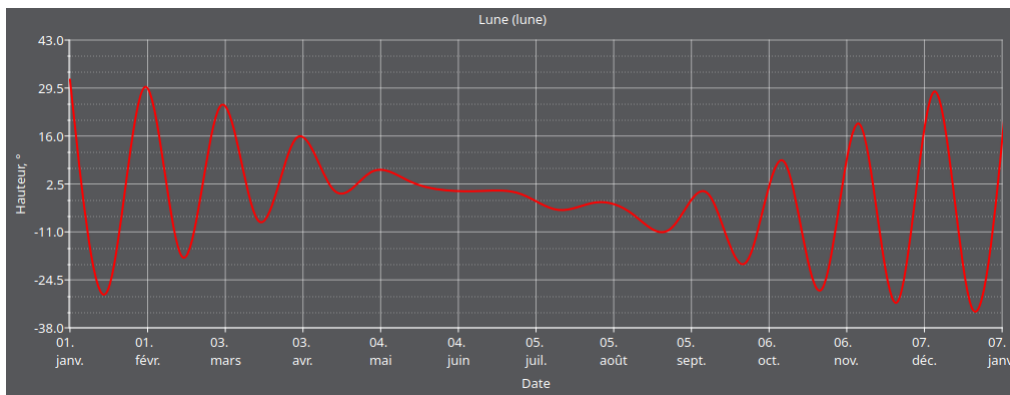


FIG. 3. Variations in the height of the Moon during 2024

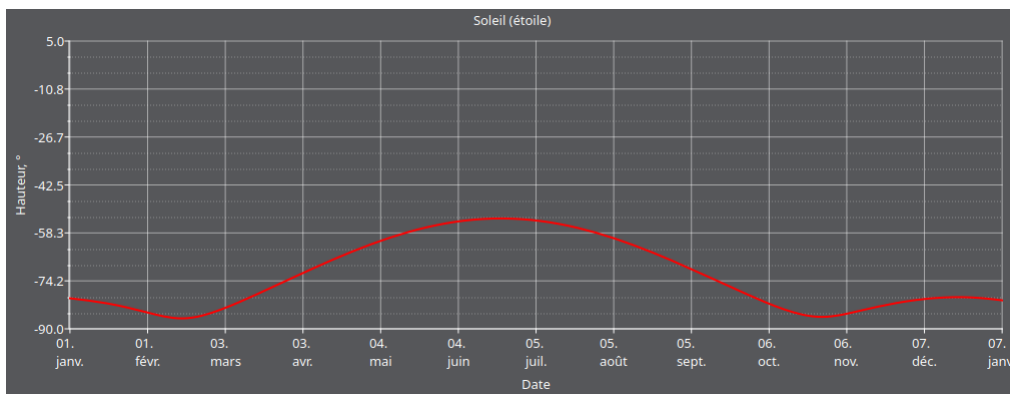


FIG. 4. Variations in the height of the Sun during 2024

These tidal phenomena elucidate the movements of the Earth’s fluids, which play a foundational role in the climate of planet Earth. The following figures illustrate the displacement of the disruptive star [8].

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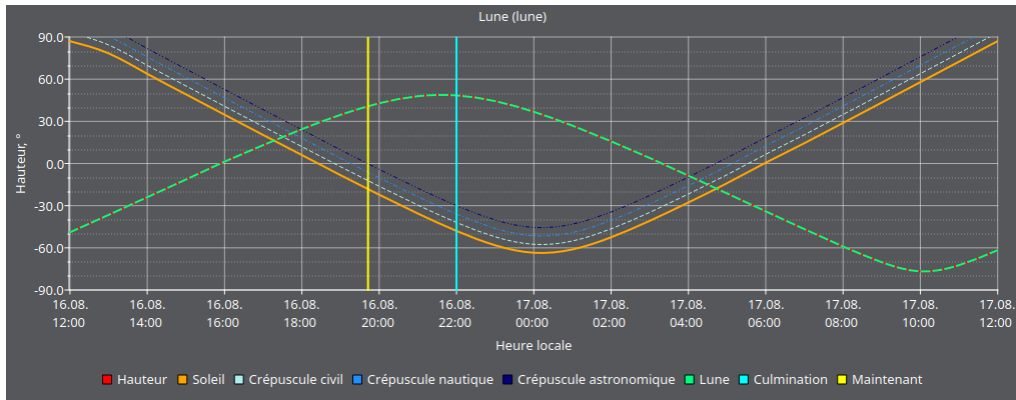


FIG. 5. Variations in the height of the Moon and the Sun on 23/10/2024

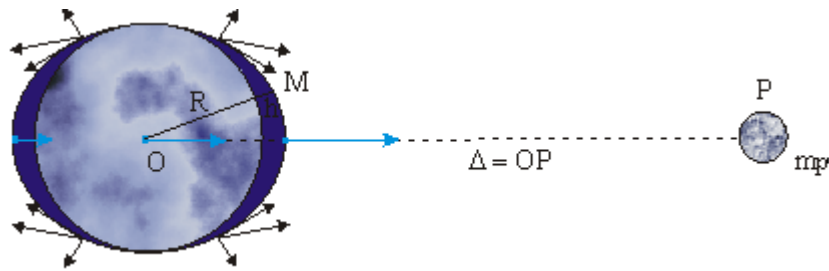


FIG. 6. The disruptive body in the earth equator

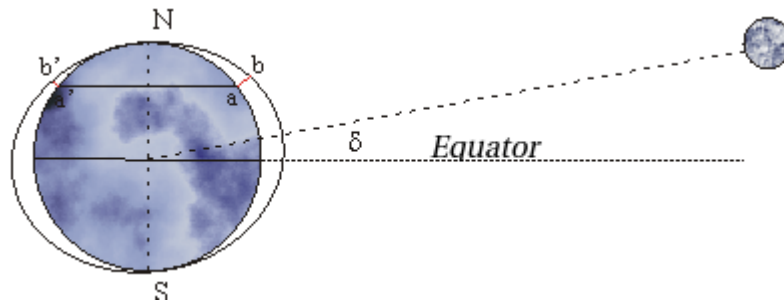


FIG. 7. The perturber body is not on the Earth's equator

The tidal height is given by the following formula as a function of the star i (Sun, Moon, etc.):

$$h_{fluidi} = \frac{Gm_{pi}R^2}{2g\Delta} (3\cos^2\chi_i - 1) \tag{13}$$

$$h_{fluidT} = \sum_i \frac{Gm_{pi}R^2}{2g\Delta} \{3\cos^2[\cos^{-1}(\frac{\sin h_i - \sin \delta_i \cdot \sin \varphi_i}{\cos \delta_i \cdot \cos \varphi_i})] - 1\} \tag{14}$$

• h_{fluidT} : total height of the fluid tide

• g : $9,81m/s^2$

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•**G**: gravitational constant $G = 6.67259 * 10^{-11} m^3 kg^{-1} s^{-2}$

• **χ_i** : (\vec{OP}_i, \vec{OM}_i)

In these tidal movements, shear due to Kelvin-Helmholtz instabilities is generated, playing a crucial role in the speeds and forces of the harmathan and the monsoon. Kelvin-Helmholtz instabilities are caused by a shear phenomenon which can occur in a single fluid (due to a velocity gradient in the normal direction). Such instabilities may occur in a single fluid (due to a velocity gradient in the direction normal to that of the flow), in a single fluid (due to a velocity gradient in the direction normal to that of the flow), or at the interface between two superimposed fluids moving at different speeds. These instabilities manifest as self-coiling vortices [9]. They can, in rare instances, be observed at the summit of specific clouds on Earth, designated as fluctus clouds. Additionally, Kelvin-Helmholtz instabilities manifest in the upper atmosphere of gas giants, where they are driven by the disparity in velocity between the jet streams that form parallel bands to the equator of these planets. We may consider a problem in which two layers of identical height, h , comprising immiscible fluids, are superimposed and separated by an interface in the presence of surface tension and in a gravitational field. The theoretical study of the Kelvin-Helmholtz instability is conducted through a linear stability analysis of the problem, which enables the determination of a dispersion relation. Subsequently, the dispersion relation is employed to ascertain the requisite condition for a wave number k to be unstable [10].

$$(u_1 - u_2)^2 > \frac{\rho_1 + \rho_2}{\rho_1 \rho_2} [\gamma k + (\rho_1 - \rho_2) \frac{g}{k}] \tanh(kh_{fluid} T) \tag{15}$$

$$(u_1 - u_2)^2 > \sum_i \frac{\rho_1 + \rho_2}{\rho_1 \rho_2} [\gamma k + (\rho_1 - \rho_2) \frac{g}{k}] \tanh(k \frac{Gm_{pi} R^2}{2g\Delta} \{3 \cos^2 [\cos^{-1}(\frac{\sinh_i - \sin \delta_i \cdot \sin \phi_i}{\cos \delta_i \cdot \cos \phi_i})] - 1\}) \tag{16}$$

with u following this equation[?]:

$$u(r \in \{x, y, z\}) = [\lambda^2 \alpha \sum_{n=1}^{+\infty} \frac{[2(\lambda r - \frac{t}{\tau_c})]^n}{2n!} - (N - 10) \sum_{n=1}^{+\infty} \sum_{j=1}^{+\infty} \frac{(500n + 2)^{j-1}}{n! j! (500n + 1)} (\lambda r - \frac{t}{\tau_c})^j + \lambda^2 \frac{f r^2}{2} - \lambda^2 \frac{Pr}{\rho} \frac{\tau_c}{\omega^2 t} + \frac{\lambda^2 \tau_c}{t} \int \sum_{n=1}^{+\infty} \frac{(\lambda r^{1-n} - \frac{tr^{-n}}{\tau_c})^n}{n!} dr \tag{17}$$

Using indices 1 and 2 for the upper and lower layers respectively and denoting γ the surface tension coefficient and k the (angular) wavenumber, defined from the wavelength λ by $k = 2\pi/\lambda$. The case of Rayleigh-Taylor instability is found in the absence of surface tension if $\rho_1 > \rho_2$

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[11]. All the wavenumbers are then unconditionally unstable (the condition is verified whatever the value of k since the left-hand member is positive or zero and the right-hand member is strictly negative). The wavenumber function, represented by the variable k , reaches a minimum on the right-hand side of Equation (16) at the critical wavenumber, k_c . This allows the critical value of the velocity difference, u_c , to be determined, beyond which the Kelvin-Helmholtz instability can develop. In the event that the velocity difference is less than the aforementioned critical value, all wave numbers are deemed to be stable[12]. In all other cases, a range of values for k are unstable, resulting in the development of Kelvin-Helmholtz instabilities. Initially, these instabilities manifest linearly, but then undergo a rapid transition to a nonlinear state at the interface [13]. The critical wave number and the critical velocity difference are given by the following equations:

$$k_c = \sqrt{\frac{g(\rho_2 - \rho_1)}{\gamma}} \tag{18}$$

$$\Delta u_c = \left(2 \frac{\rho_1 + \rho_2}{\rho_1 \rho_2} \sqrt{g(\rho_2 - \rho_1)\gamma}\right)^{1/2} \tag{19}$$

For further theoretical details concerning Kelvin-Helmholtz instabilities, we direct the interested reader to Chandrasekhar. The experimental apparatus employed by Thorpe (see Figure 8) provides an excellent opportunity to study the development of Kelvin-Helmholtz instabilities, which occur in a sheared flow at the interface between two immiscible fluids [13]. Our main results are deduced

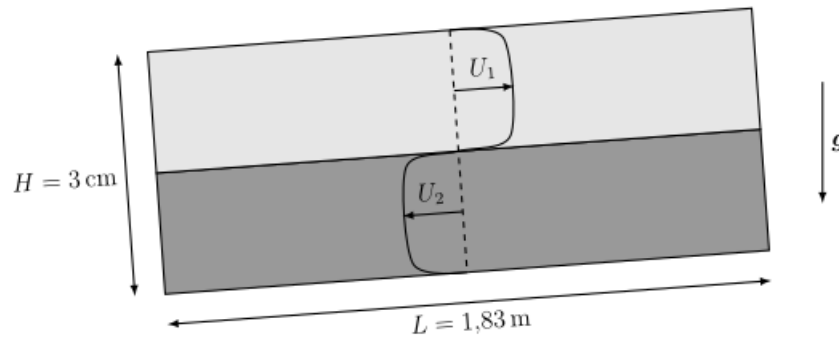


FIG. 8. Thorpe’s experiment

from climate satellite observations.

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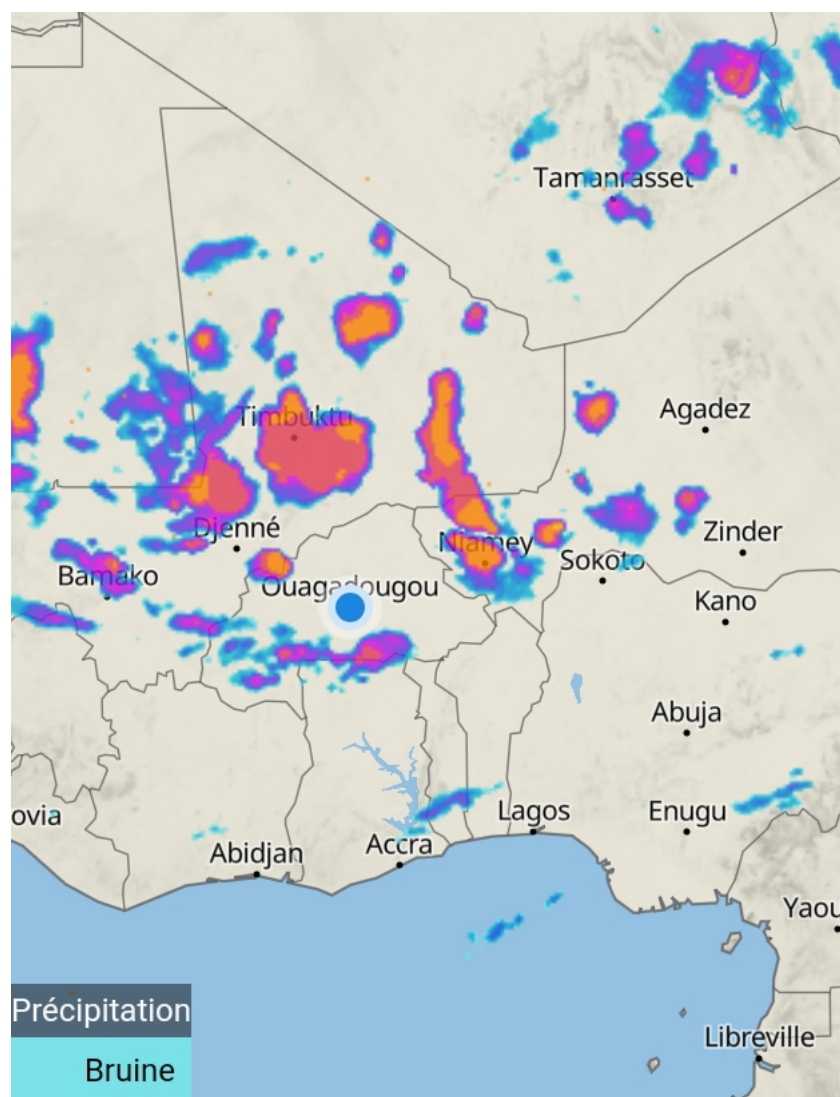


FIG. 10. Influence on oragic formations on the observed planetary domain

Further satellite observations were conducted to gain further insight into the formation of these expansive clouds. These observations demonstrate that each layer possesses a distinct density resulting from the accumulation of gases at varying heights. The consequences of this is that wind speed in a specific direction within an atmospheric layer gives rise to the circulation of adjacent layers in the opposite direction, as a result of Kelvin-Helmholtz instabilities.

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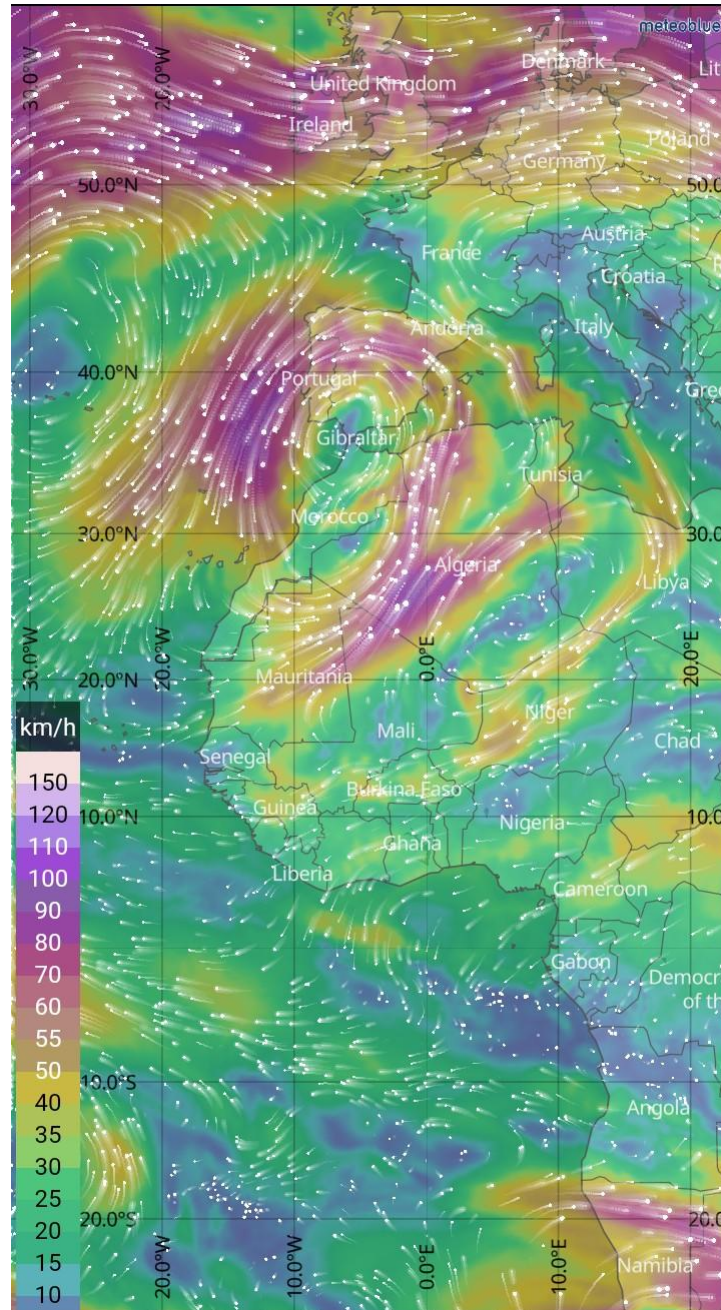


FIG. 11. Wind circulation at an altitude of 700Mo

As a consequence of the Earth's surface insolation, the planet's temperature rises, thereby acting as a heat source for the surrounding atmosphere. As the air temperature increases, it expands and, as a consequence of this expansion, its density is reduced. The circulation of cold winds, caused by atmospheric tides above the given area, results in an upward suction in the opposite direction due to Kelvin-Helmholtz instabilities of the warm air at low altitudes near the ground. This leads to the formation of a circulation. As the air warms, it is able to absorb more water,

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which in turn causes a cooling effect and the gradual disappearance of humidity. This results in the formation of clouds at higher altitudes. Noted that once reach the 700Mo altitude, they are carried by this layer winds.

It can be seen that the phenomenon of atmospheric tides generated by the heights of the moon

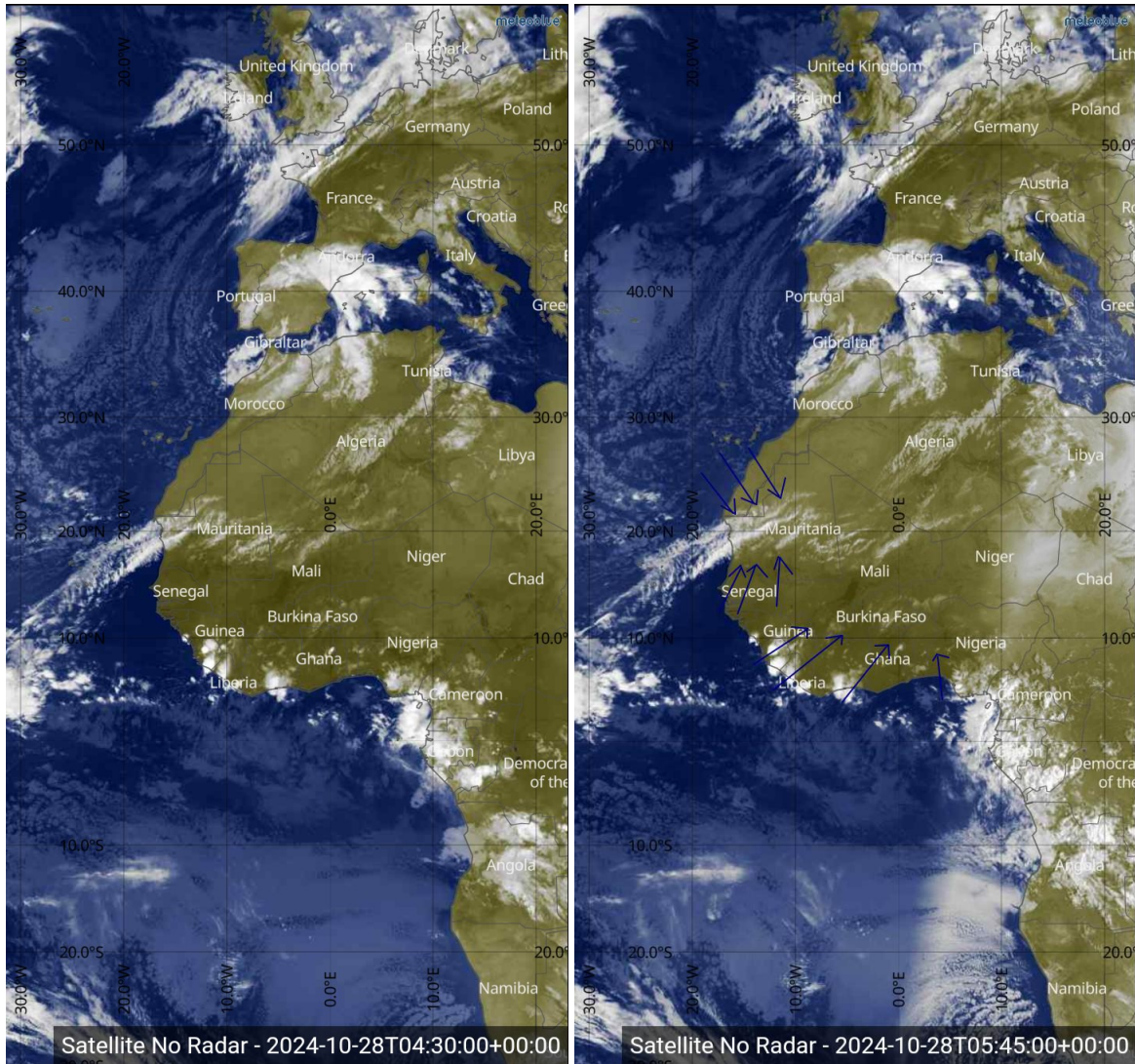


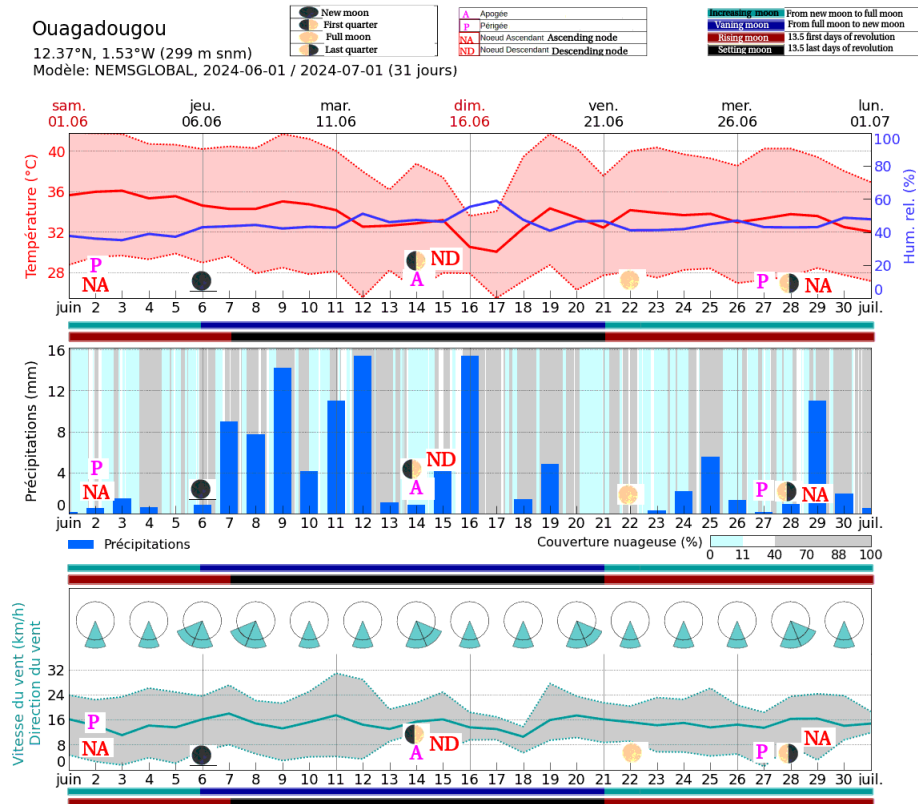
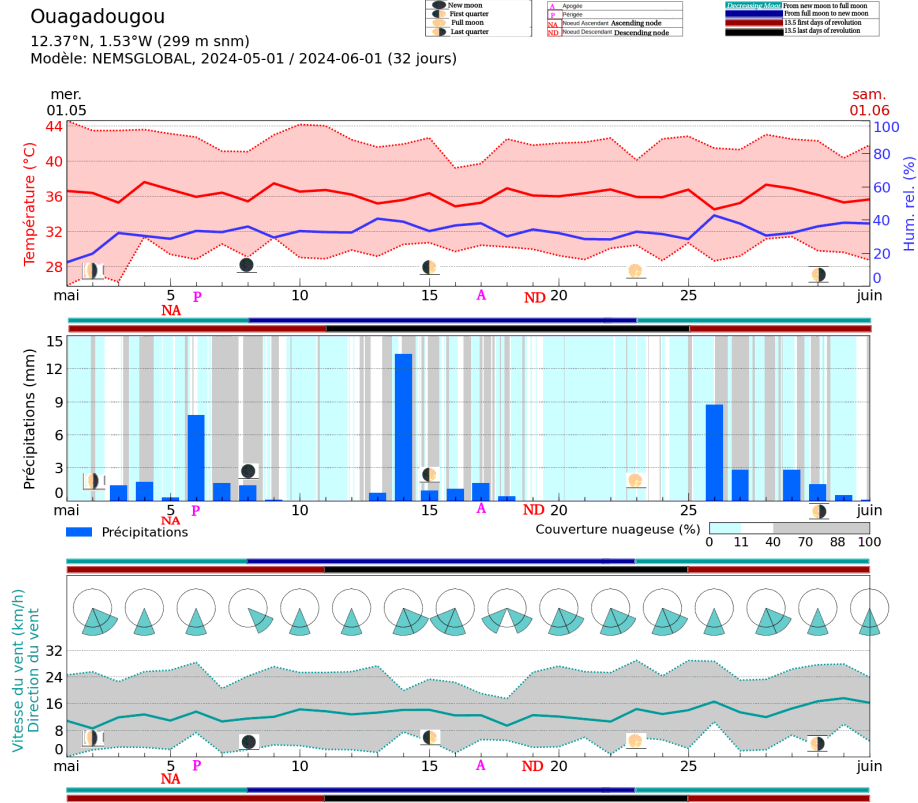
FIG. 12. Movement in the opposite direction to the winds of the lower atmospheric layers, which are charged with water vapour, occurs as the vapour condenses.

and the sun has a considerable impact on the climate, which should not be overlooked. It is imperative that these phenomena be accorded priority in the study of climatic phenomena on planet Earth. The alteration in air density resulting from the accumulation of gases produced by human activity and solar storm generated by sun activities exerts an influence on the distribution of atmospheric high tides across the planet, which in turn affects the distribution of winds, temperatures

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and precipitation. The following figures illustrate the distribution of temperatures, precipitation and winds in Ouagadougou as a function of the heights of the Moon and Sun for a period from May to September 2024.

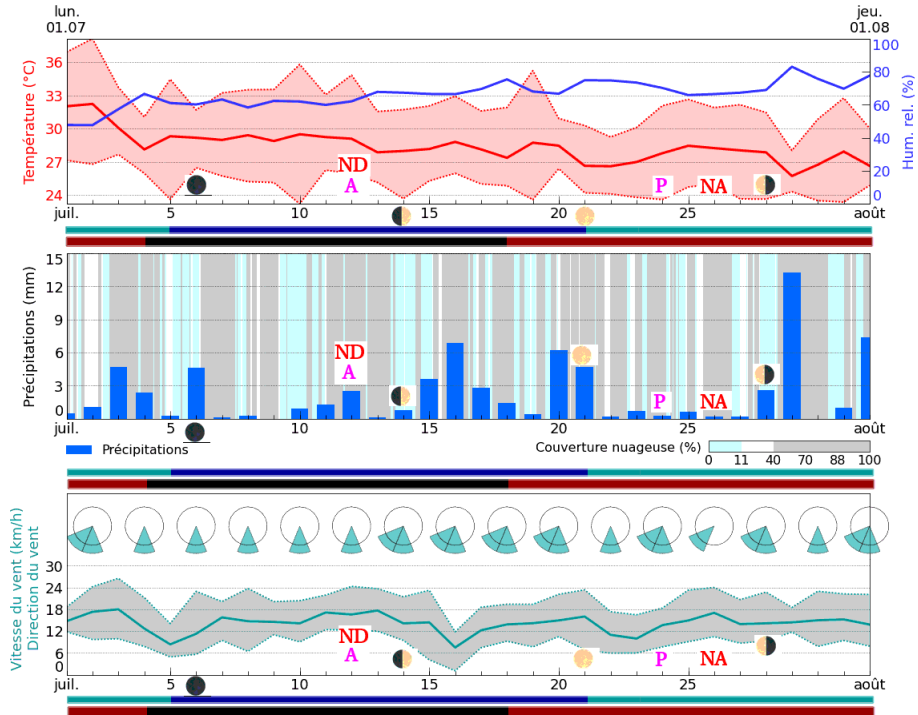
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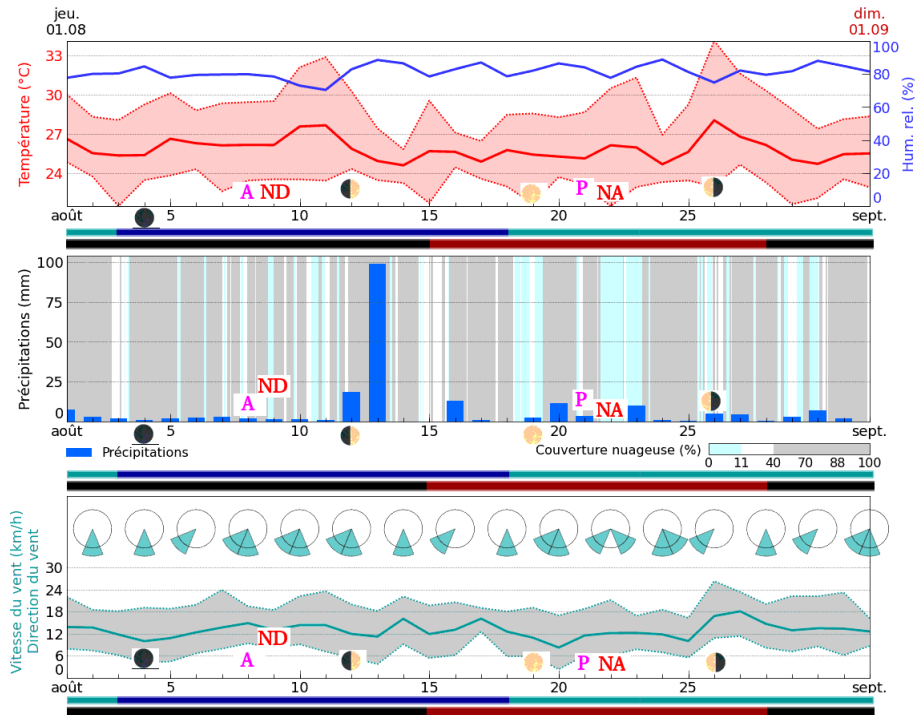
Ouagadougou

12.37°N, 1.53°W (299 m snm)
 Modèle: NEMSGLOBAL, 2024-07-01 / 2024-08-01 (32 jours)



Ouagadougou

12.37°N, 1.53°W (299 m snm)
 Modèle: NEMSGLOBAL, 2024-08-01 / 2024-09-01 (32 jours)



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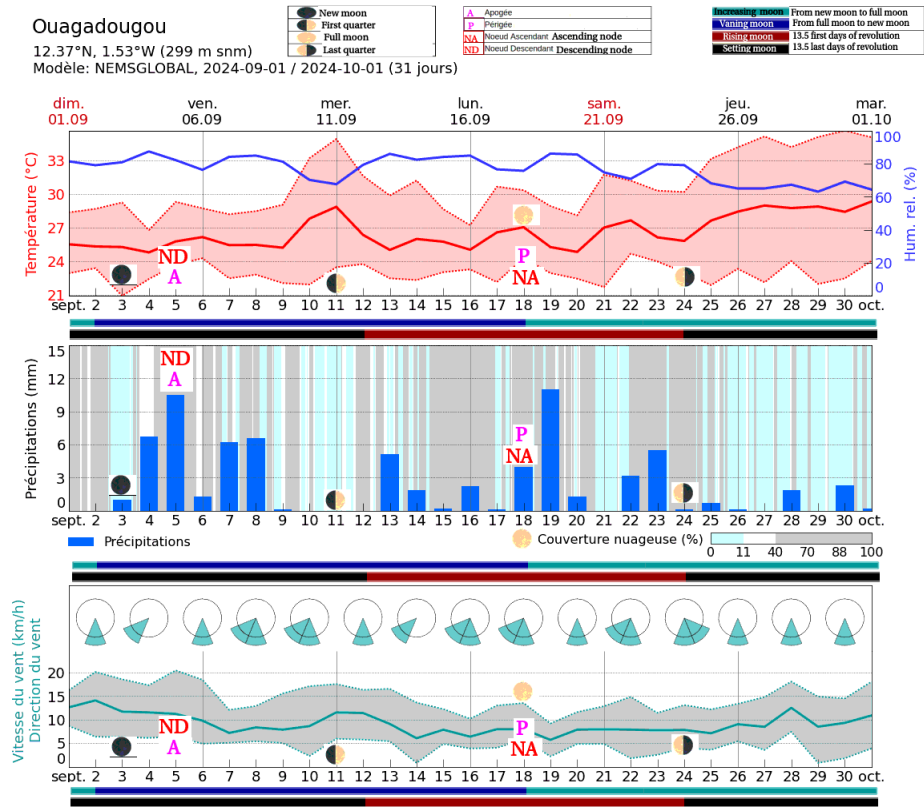


FIG. 13. Movement in the opposite direction to the winds of the lower atmospheric layers, which are charged with water vapour, occurs as the vapour condenses.

Specific case study and zone rainfall function

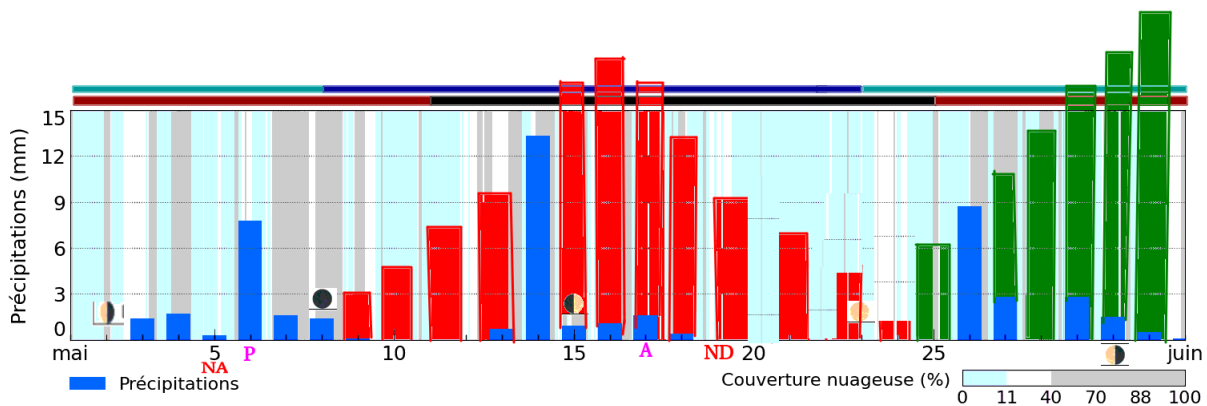


FIG. 14. Gaussian rain distribution

Figure 14 shows blue bars for rainfall obtained during the months, red and green bars that must be used. To explain this, suppose that during the rainy month in the zone, four (4)

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Gaussian rainfall functions must be taing into account and over the months studied:

$$P_i = \frac{P_{imax}}{\sigma_i\sqrt{2\pi}} \exp\left[-\frac{1}{2}\left(\frac{x-\mu_i}{\sigma_i}\right)^2\right] \cos^2(\phi_i) \quad (20)$$

- $\cos^2\phi_i$: Giving the daily water distribution during the Moon increasing, decreasing, rising setting activities period.
- $\frac{P_{imax}}{\sigma_i\sqrt{2\pi}} \exp\left[-\frac{1}{2}\left(\frac{x-\mu_i}{\sigma_i}\right)^2\right]$: Moon activity period gaussian.

99 percent of Moon activity period is in $[x - 3\sigma_i, x + 3\sigma_i]$. To get a best rainfall prediction, we need to take into- account the Moon activity days interval corresponding to increasing, decreasing, rising and setting Moon. The general prediction is define as

$$P_M = \frac{P_{I\max}}{\sigma_I\sqrt{2\pi}} \exp\left[-\frac{1}{2}\left(\frac{x-\mu_I}{\sigma_I}\right)^2\right] \cos^2(\phi_I) - \frac{P_{S\max}}{\sigma_S\sqrt{2\pi}} \exp\left[-\frac{1}{2}\left(\frac{x-\mu_S}{\sigma_S}\right)^2\right] \cos^2(\phi_S) \\ - \frac{P_{D\max}}{\sigma_D\sqrt{2\pi}} \exp\left[-\frac{1}{2}\left(\frac{x-\mu_D}{\sigma_D}\right)^2\right] \cos^2(\phi_D) + \frac{P_{R\max}}{\sigma_R\sqrt{2\pi}} \exp\left[-\frac{1}{2}\left(\frac{x-\mu_R}{\sigma_R}\right)^2\right] \cos^2(\phi_R) \quad (21)$$

The **TOUGMA's climate model** is summarize as follow:

$$\left\{ \begin{array}{l} u(r \in \{x, y, z\}) = \left[\lambda^2 \alpha \sum_{n=1}^{+\infty} \frac{[2(\lambda r - \frac{t}{\tau_c})]^n}{2n!} - (N-10) \sum_{n=1}^{+\infty} \sum_{j=1}^{+\infty} \frac{(500n+2)^{j-1}}{n!j!(500n+1)} (\lambda r - \frac{t}{\tau_c})^j + \lambda^2 \frac{fr^2}{2} - \lambda^2 \frac{Pr}{\rho} \right] \frac{\tau_c}{\omega^2 t} \\ + \frac{\lambda^2 \tau_c}{t} \int \sum_{n=1}^{+\infty} \frac{(\lambda r^{1-n} - \frac{tr^{-n}}{\tau_c})^n}{n!} dr \\ (u_1 - u_2)^2 > \sum_i \frac{\rho_1 + \rho_2}{\rho_1 \rho_2} [\gamma k + (\rho_1 - \rho_2) \frac{g}{k}] \tanh(k \frac{Gm_{pi} R^2}{2g\Delta} \{3 \cos^2[\cos^{-1}(\frac{\sinh_i - \sin\delta_i \cdot \sin\phi_i}{\cos\delta_i \cdot \cos\phi_i})] - 1\}) \\ P_M = \frac{P_{I\max}}{\sigma_I\sqrt{2\pi}} \exp\left[-\frac{1}{2}\left(\frac{x-\mu_I}{\sigma_I}\right)^2\right] \cos^2(\phi_I) - \frac{P_{S\max}}{\sigma_S\sqrt{2\pi}} \exp\left[-\frac{1}{2}\left(\frac{x-\mu_S}{\sigma_S}\right)^2\right] \cos^2(\phi_S) - \frac{P_{D\max}}{\sigma_D\sqrt{2\pi}} \exp\left[-\frac{1}{2}\left(\frac{x-\mu_D}{\sigma_D}\right)^2\right] \cos^2(\phi_D) \\ + \frac{P_{R\max}}{\sigma_R\sqrt{2\pi}} \exp\left[-\frac{1}{2}\left(\frac{x-\mu_R}{\sigma_R}\right)^2\right] \cos^2(\phi_R) \end{array} \right.$$

SCOPE OF FUTURE WORK

The main climate change is the atmospheric tides due to Sun and Moon height. However, greenhouse gases have an impact on atmospheric layers density and the distribution and volume of these atmospheric tides. This work encourages a deeper comprehension of the dispersion of elevated atmospheric tides in both time and space, with the objective of developing more accurate forecasts and, consequently, the prevention of potential climatic catastrophes that could have adverse effects on populations. A deeper comprehension of this distribution will facilitate a more accurate assessment of the actual impact of anthropogenic greenhouse gases on the climate. This,

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in turn, will prove invaluable in formulating effective climate change mitigation strategies that will benefit humanity.

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